

LEARNING RESOURCE MEASURING THE SUN

SCIENCE Years 7–10
EARTH AND SPACE Stage 5

WHAT IS THE AIM:

To calculate the diameter of the Sun in km.

EXPERIMENT:

We will record the time the Sun takes to cross a line as seen through a telescope. This is called the transit time, T .

If we know the distance to the Sun d and the rate of the Earth's rotation, basic trigonometry can be used to calculate the diameter of the Sun, D in km.

THEORY:

As the Earth rotates, the Sun appears to circle the Earth (360°) once each day (24 hours).

Half of the apparent angular size, θ° , of the Sun, α° * is equal to a fraction of that 360° circle.

* Why will we use half of the angular diameter of the Sun and half of the transit time?

We know half of the transit time, t is a fraction of a day and half of the apparent Solar diameter, α is a fraction of a full circle (note the units).

That is:
$$\frac{\alpha^\circ}{360^\circ} = \frac{t_{\text{sec}}}{24_{\text{h}}}$$
 Eqn (i)

Therefore
$$\alpha = \frac{\boxed{} \times \boxed{}}{\boxed{}}$$

However, our time will be in seconds, so convert 24 hours to seconds:

$$\alpha = \frac{\boxed{} \times t}{\boxed{} \times \boxed{} \times \boxed{}} = \frac{\boxed{} \times t}{\boxed{}} = \frac{t}{\boxed{}}$$

$\alpha = \underline{\hspace{2cm}}$

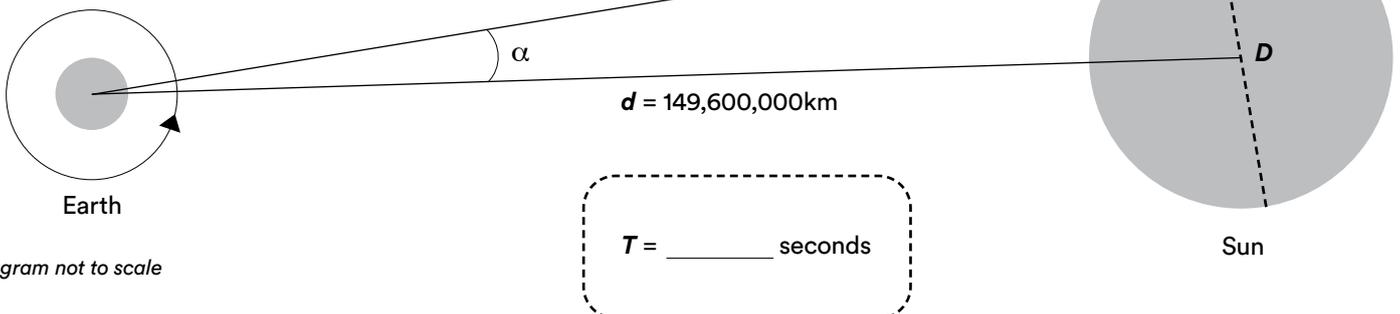


Diagram not to scale

If we now use the Sine ratio and look at the diagram on the far right we see the following:

$$\sin(\alpha) = \frac{\boxed{}}{\boxed{}} = \frac{\boxed{}}{\boxed{}} \quad \text{Eqn (ii)}$$

Rearranging equation (ii), we now calculate the solar radius

$$r = \sin\left(\frac{\boxed{}}{\boxed{}}\right) \times \boxed{}$$

$r = \underline{\hspace{2cm}}$ km, and finally

$$D = 2 \times r \quad \text{Eqn (iii)}$$

$D = \underline{\hspace{2cm}}$ km

ADDED COMPLICATIONS:

The value of d (shown below) is an average, from centre to centre. If you make your own transit device to record the time you will also need to find the Earth-Sun semi-major axis distance for that day.